

RENSSELAER OBSERVATORY PUBLICATIONS

Number 15

18-Megacycle Cosmic-Noise Intensities

1959 December 1 to 1960 January 31

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INTRODUCTION

The Data

Apparatus for measuring the changes in intensity of 18-megacycle cosmic noise received from the sky has been set up at the Sampson Station of the Observatory of Rensselaer Polytechnic Institute. The position of the equipment is latitude $42^{\circ} 47'$ North, longitude $73^{\circ} 27'$ West. The purpose of the program is to detect the occurrence of solar flares indirectly by the associated effect on the transmission properties of the ionosphere. The effect is to decrease suddenly the transmissivity of the ionosphere to produce a sudden cosmic-noise absorption (SCNA), examples of which may be seen on the succeeding pages of this paper. The apparatus responds in a somewhat similar fashion to several other effects, particularly auroral activity. At times of solar disturbance there are in addition a number of increases of incoming radiation which we believe to be solar in origin. In what follows these are referred to as bursts.

This is a continuation of the records of 18-megacycle cosmic-noise intensity published in ROP 1. The same receiver is in use. The antenna, however, was rebuilt on June 19, 1958 to the same characteristics as the former one. The new antenna consists of tubular twin-lead instead of the former open lines, and its performance is presumably more independent of moisture conditions.

On the pages which follow we show reproductions of the recorder tapes, one week on a page. The original tapes move at four inches per hour. The reduction for the photographs here is very closely 1 to 12. Times are given in Universal Time.

It should be noted that the receiver response is rapid when the signal decreases, but slow when an increase occurs. Thus the record of a burst gives only an approximation to the rate of rise and maximum value attained, but generally not the true values. The times of beginning and the duration, however, are significant.

The vertical markers every ten minutes are produced by a separate receiver tuned to station WWV at 5 megacycles. The amplitude of these markers during a solar disturbance is very rough indication of the field strength of the 5-megacycle signal as received at the Sampson Station.

activity, on the other hand, shows a slower decrease of intensity, and usually a more rapid rise, resulting in a more symmetrical curve. Such events are often repeated a number of times during a single night, with varying values of maximum absorption. There are occasionally isolated daytime events in which the decrease in intensity is slow; we refer to these as "slow SCNA's."

The bursts generally show a fast onset, limited by the time constant of the receiver for increasing signals, and a short duration. Because of the time constant, the maximum intensity recorded may be considerably less than the true maximum intensity. The time of beginning and the duration are, however, significant. Often the bursts occur in groups. Noise storms on the sun may show on the records as a series of superimposed bursts, giving a highly variable trace.

Calibration

On most days two two-point calibrations are recorded, usually at about 0100 and 1100 U.T., for the purpose of checking receiver stability. The lower, "cold," step is produced by substituting a cold resistor for the antenna and indicates the receiver noise level. The upper, "hot," step is an arbitrary level approximately equal to the maximum level of incoming cosmic noise. In addition, a daily step calibration shows 3-decibel steps ($V_n^2/V_{n+1}^2 = 0.5$) from the same "hot" level, called 0 db. The steps are produced by feeding known voltages from a signal generator into the receiver. The lowest step is the "cold" calibration rather than a known input signal. The changes from day to day are small, but sufficient to make the scale at the left of the charts only approximate. Note the non-linearity of the scale.

The Comments

On the pages facing the reproductions are lists of events and comments about the records. In the comments, certain numerical indications of importance and intensity are given. For SCNA's the "class" is a measure of importance on a rising scale of 1- to 3+, determined by amplitude and duration. The percentage of absorption is the intensity ratio of the least cosmic noise received during the event to the noise which would have been received if the event had not occurred.

In some cases, a rise or fall of the recorder pen can be identified with a disturbance caused by operation or testing of other equipment in or near the building in which the 18-megacycle receiver is located. Such a rise or fall is identified as "interference." In some cases it is suspected rather than established. "b" means before.

Beginning January 1, 1960 the comments will include the Geophysical Alerts and Special World Intervals which are issued by the World Warning Agency of the International World Day Service.

Acknowledgements

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Justin A. Curtis

Ralph Haskell

Masakazu Oshima

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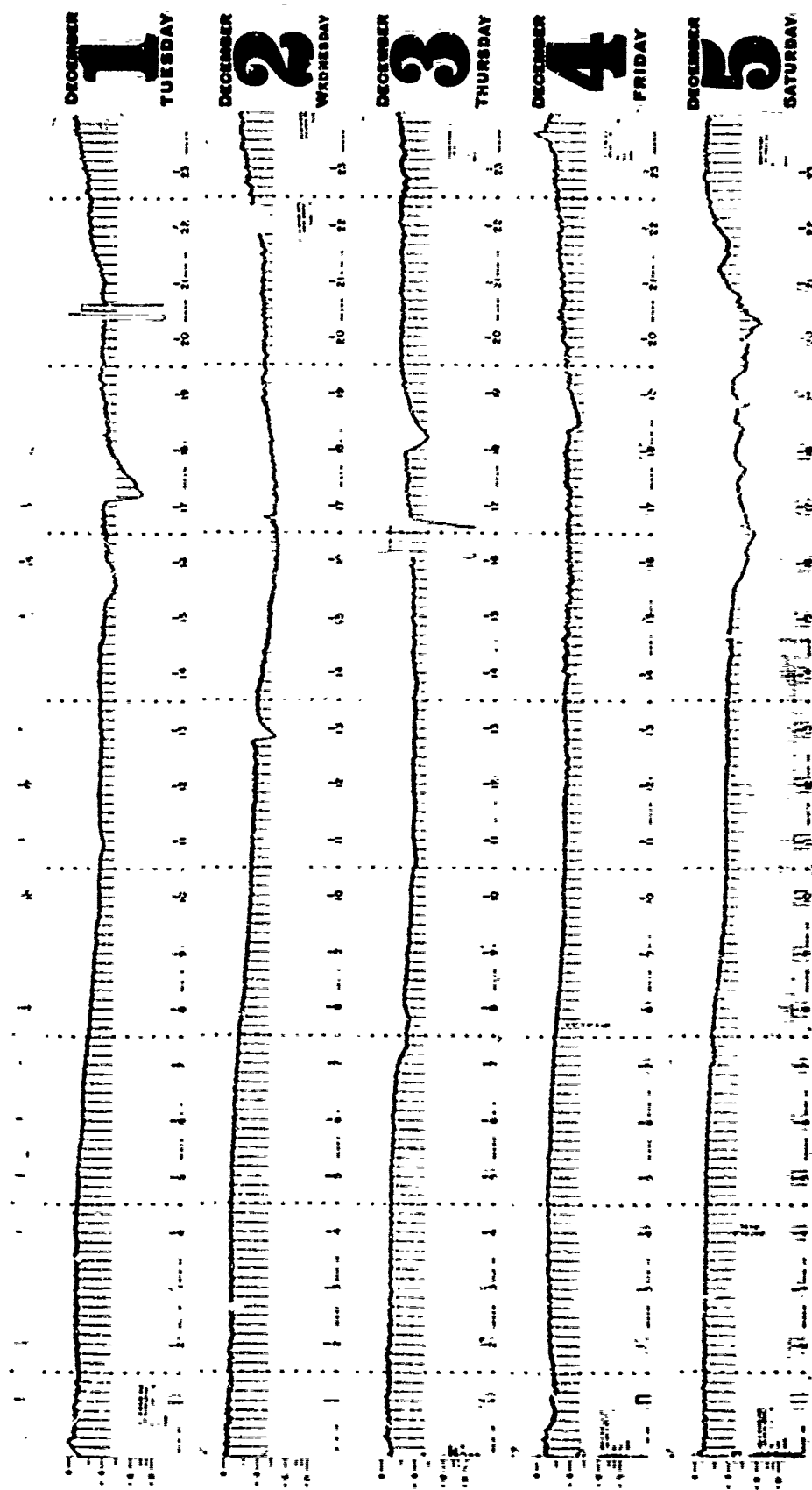
Erratum

ROP 14, page 8, November 21, 1959
Insert time, 1620-1720, in front of comment.

ROP 15, 1960

1959

December 1	1705 - 1820	SCNA, class 3, 69% absorption.
December 2	1248 - 1318 1644 - 1649	SCNA, class 2, 49% absorption. Burst.
December 3	1756 - 1910	SCNA, class 2, 49% absorption.
December 4	1820 - 1906	SCNA, class 1, 30% absorption.
December 5	1530 - 2206	Absorption, possibly caused by daytime aurora, during IGY alert. 54% maximum absorption.

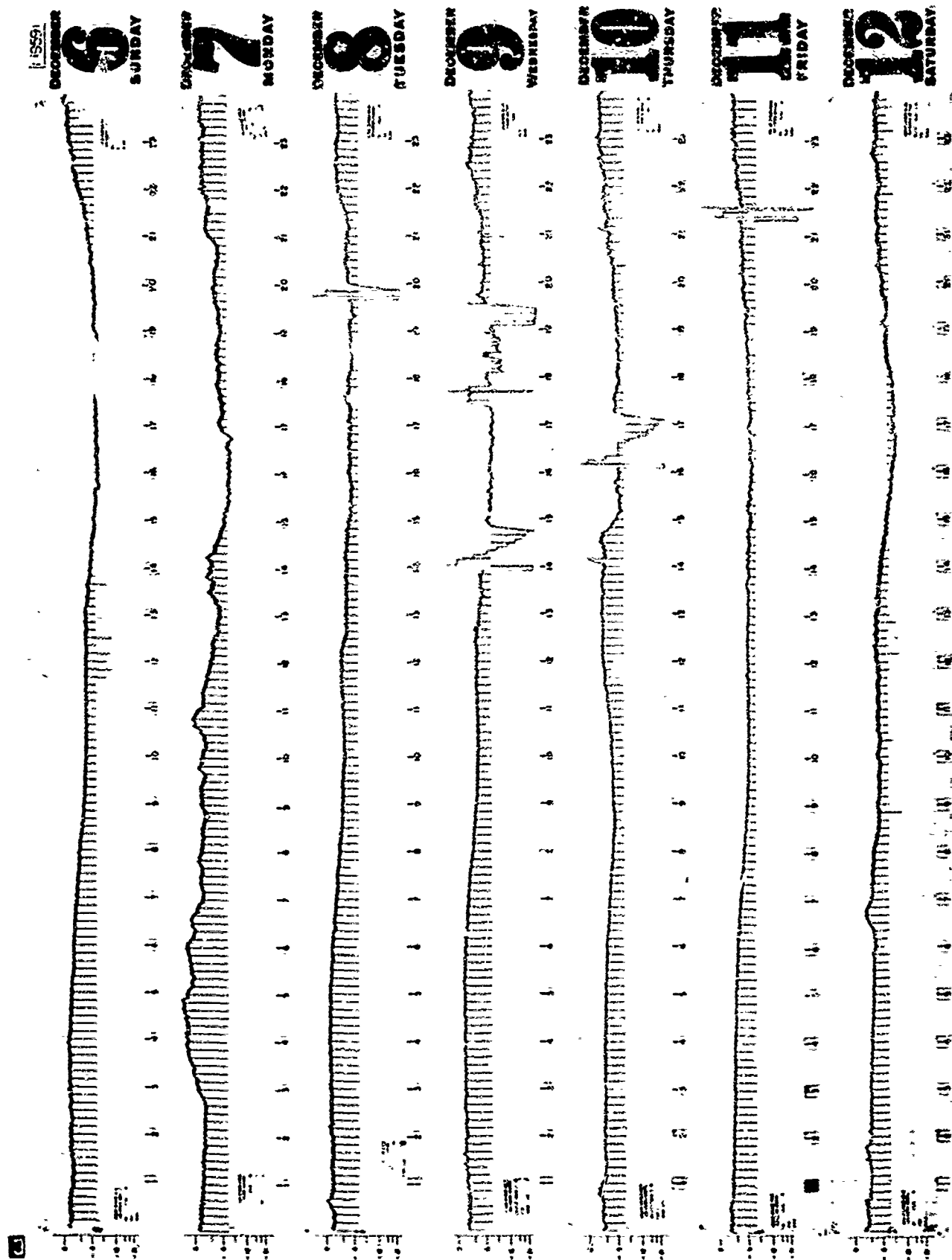


18 MC COSMIC NOISE INTENSITY VS. UNIVERSAL TIME IN HOURS

1959		
December 6	No comments.	
December 7	0242 - 2115	Variability correlates with precipitation and electric activity.
December 8	No comments.	
December 9	1722 - 1932	Equipment adjustments.
	2120 - 2400	Rise in level due to failure of heating equipment.
December 10	0000 - 1444	Unusually high level due to failure of heating equipment.
	1404 - 1410	Burst.
	1444 - 1510	Fall in level due to gradual warming up of equipment.
	2104 - 2110	Burst.
December 11	2135	Recorder zero point adjustment.
December 12	0613 - 0710	Variation in level correlates with atmospheric electricity.

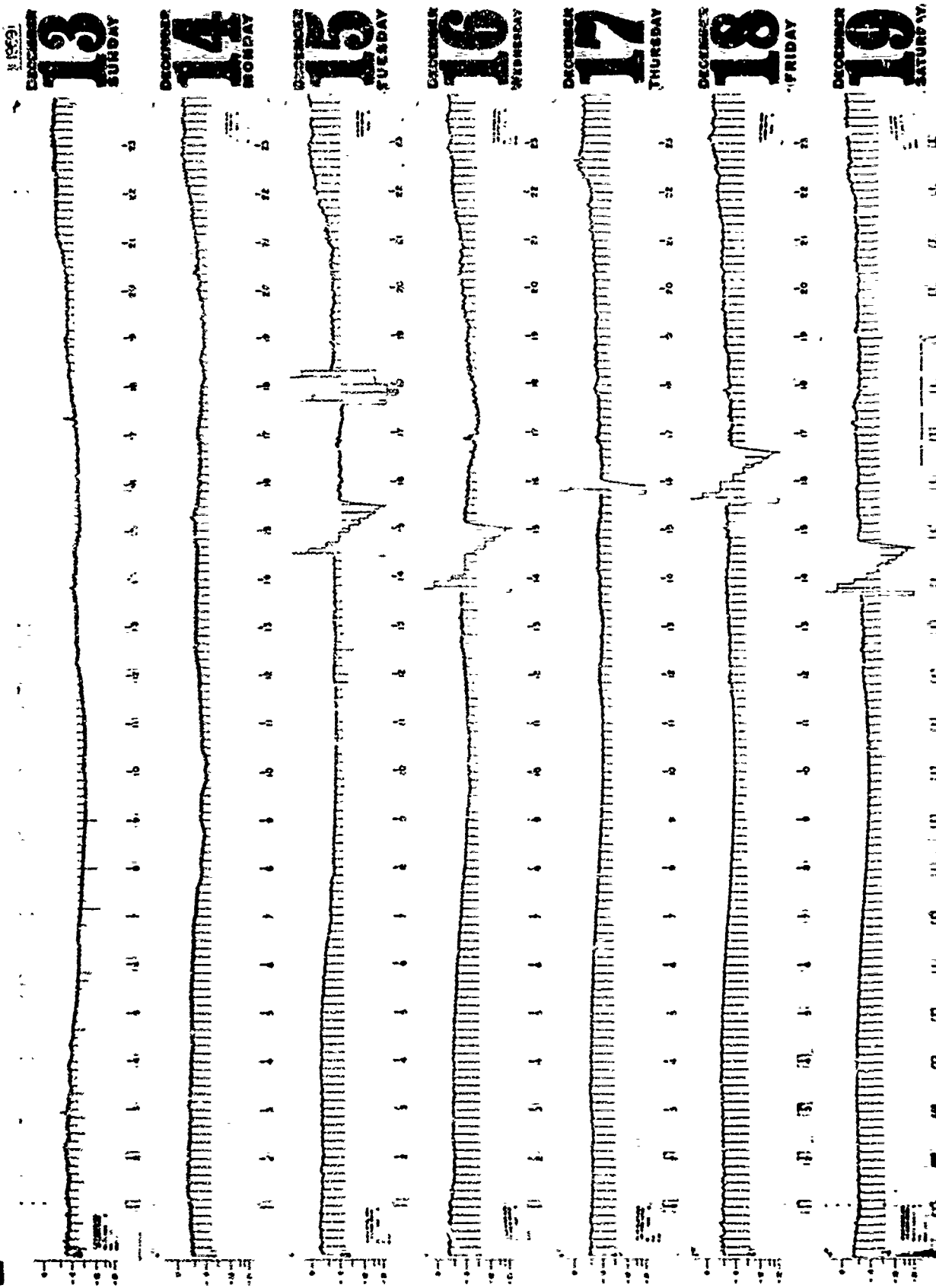
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1959

December 13	0250 - 0258 1728 - 1733	Peak, cause unknown. Burst.
December 14	No comments.	
December 15	1753 2100 2110	Recorder zero adjustment preceded and followed by hot-cold calibration. Burst. Burst.
December 16	1651 - 1659	Burst.
December 17	2215 - 2245	Bursts.
December 18	1450 - 1510 1650 - 1710 1750 - 1757	Bursts. Bursts. Burst.
December 19	1710 - 1717	Burst.



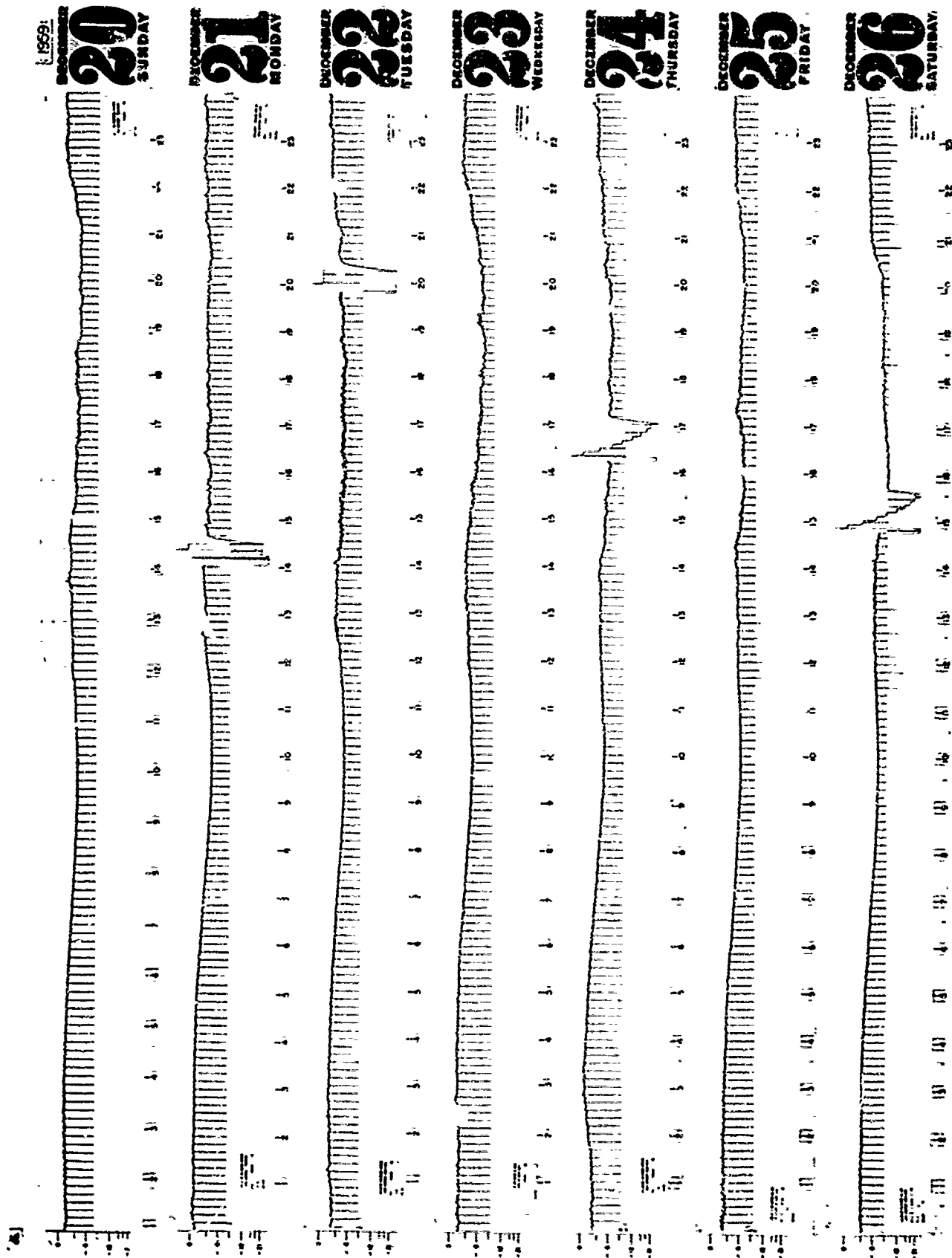
COSMIC NOISE INTENSITY VS UNIVERSAL TIME IN HOURS

ROP 15, 1960

1959

December 20 Note that the trace for this day begins at 0100 UT, not 0000 UT as usual. Some ten-minute marks are farther apart than normal.

1344 - 1350	Bursts.
December 21	
1230 - 1238	Burst.
1450 - 1730	Variable.
December 22	
1358 - 1852	Variable.
December 23	
1759 - 1810	Bursts.
1905 - 1920	Bursts.
December 24	
2007 - 2400	Variable.
December 25	No comment.
December 26	No comment.



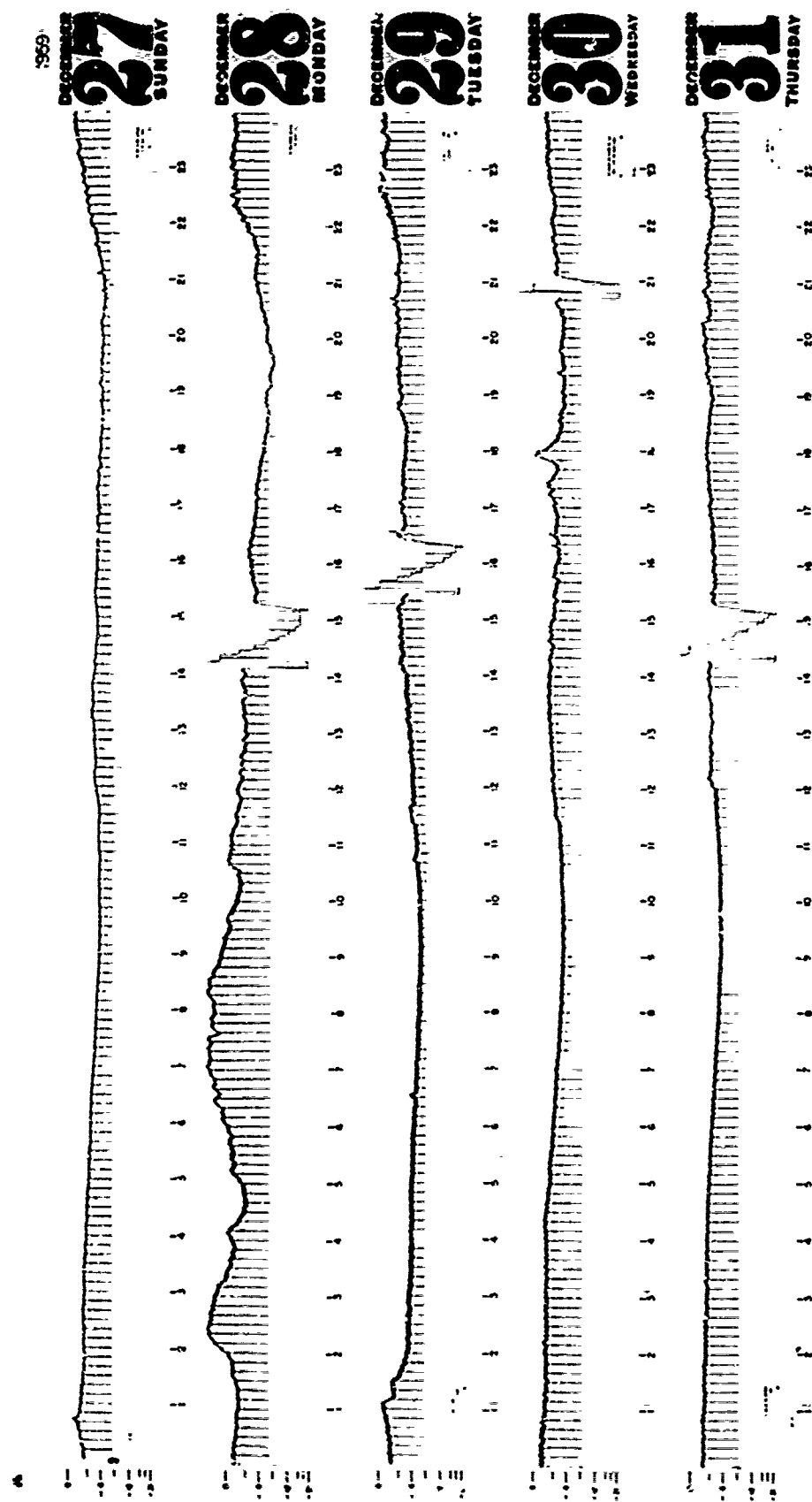
18 MC COSMIC NOISE INTENSITY VS UNIVERSAL TIME IN HOURS

1959

December 27 2047 - 2140 Bursts.

December 28 Large magnitude of trace variation possibly due to area storm. Freezing rain or rain all day.

December 29 Icing on all equipment might prevent indications of small phenomena.
 0100 - 0200 Unusually rapid drop may be continuation of variability of December 28.
 0629 - 0635 Peak, cause unknown; correlates with change in trace on most other equipment.
 1517 Equipment adjustment.
 1631 - 1635 Burst.
 December 30 1433 - 1650 Variable.
 1655 - 1820 Sudden change in level, possibly caused by heavy winds.
 December 31 2016 - 2050 Possible SCNA, class 1-, absorption 20%, recovery unusually slow.



18 MC COSMIC NOISE INTENSITY VS UNIVERSAL TIME IN HOURS

1960

January 1

1230 - 2200

Variability, cause unknown.

January 2

Note that trace for this day begins at 0100 UT. Some ten-minute marks are farther apart than usual.

1422 - 1429

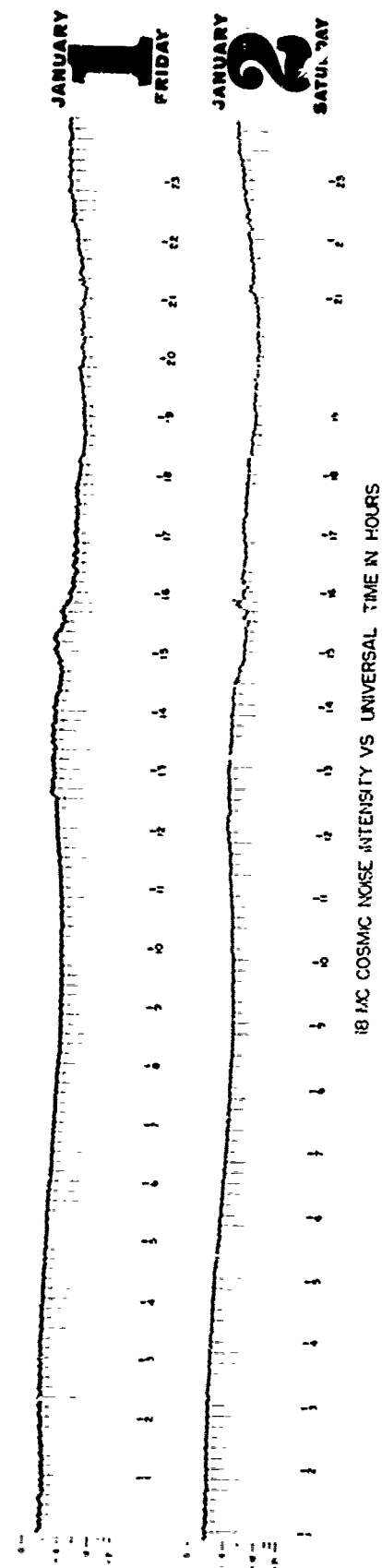
Equipment adjustment.

1525 - 1612

Several bursts.

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1960

January 3

Note that trace for this day begins at 0100 UT, not at 0000 UT as usual. Some ten-minute marks are farther apart than usual.

1217 - 1810 Variable, correlates with strong oscillation of atmospheric pressure and atmospheric electrical potential gradient.

1303 - 1317 Burst.

1325 - 1329 Burst.

1350 - 1358 Burst.

January 4

2103 - 2120 Large double burst during period of variability.

January 5

1545 - 1601 Burst.

1758 - 1803 Burst.

2057 - 2108 Double peak, correlates with precipitation.

January 6

1915 - 1945 Several large bursts.

2223 - 2400 Low trace probably due to malfunction of recorder.

January 7

0000 - 0220 Low trace due to malfunction of recorder.

2045 - 2101 Peaks, possibly due to interference.

2224 - 2233 Peak, possibly due to interference.

January 8

0115 - 1400 Trace level low and variable, probably due to malfunction of recorder.

1400 - 1711 Equipment adjustments.

1911 - 1920 Peak correlates with precipitation.

2000 - 2007 Peak correlates with precipitation.

2038 - 2045 Peak correlates with precipitation.

2150 Recorder not functioning properly.

January 9

0000 - 1410 Recorder not functioning properly.

1410 - 1655 Equipment adjustments.

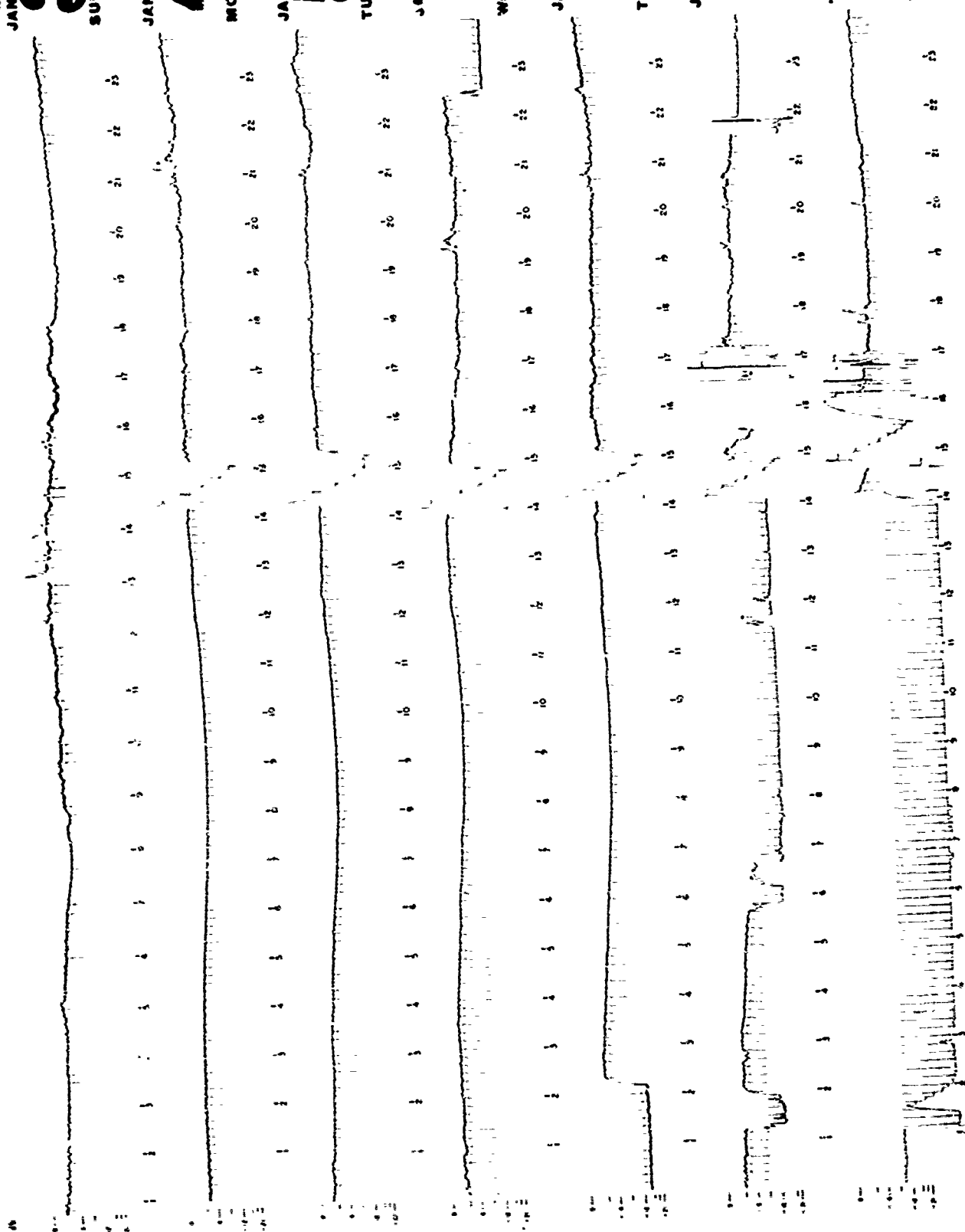
1732 - 1752 Peaks, probably due to interference or malfunction of recorder.

Beginning January 6 at 2220, the trace is probably invalid and the data are of low weight.

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1960
JANUARY 3 SUNDAY
JANUARY 4 MONDAY
JANUARY 5 TUESDAY
JANUARY 6 WEDNESDAY
JANUARY 7 THURSDAY
JANUARY 8 FRIDAY
JANUARY 9 SATURDAY

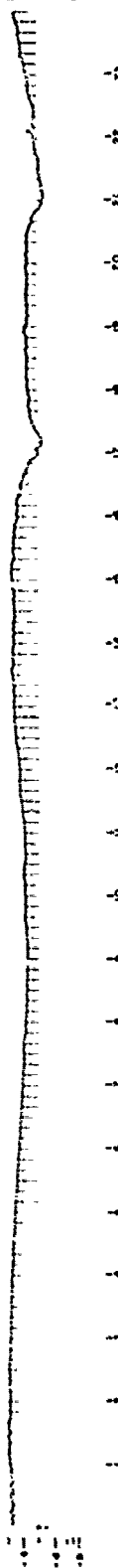


IB MC COSMIC NOISE INTENSITY VS UNIVERSAL TIME IN HOURS

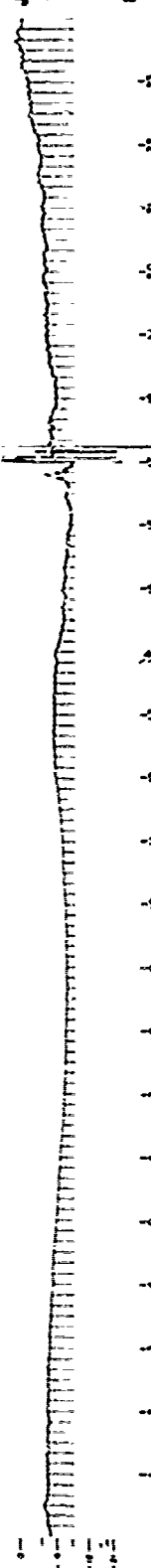
1960

January 10	1620 - 1800 2020 - 2250 2205 - 2210	Possible slow SCNA, class 3, 69% absorption. Possible slow SCNA, class 2, 58% absorptic .. Burst.
January 11	Geophysical Alert 1641 - 1655 1701 - 1715 1655 - 1722	Number 44: Magnetic storm started January 10 at 0719. Possible burst. Gain and zero-point adjustment. Gain and zero-point adjustment.
January 12	1800 - 1807 1915 - 2000	Possible burst. Variable.
January 13	Geophysical Alert 1825 - 2400	Number 45: Magnetic storm started January 13 at 1900. Variable.
January 14	0000 - 0850 1220 - 1330 1600 - 2215	Variable. Rise in level probably due to precipitation. Variable, correlates with pressure and atmospheric electric potential gradient fluctuations. Also, precipitation was reported from 1800 - 2000.
January 15	0122 - 0345 1712 - 1731 1757 - 1805	Variable, cause unknown. Peaks, cause unknown. Peaks, cause unknown.

1960
JANUARY
10
SUNDAY



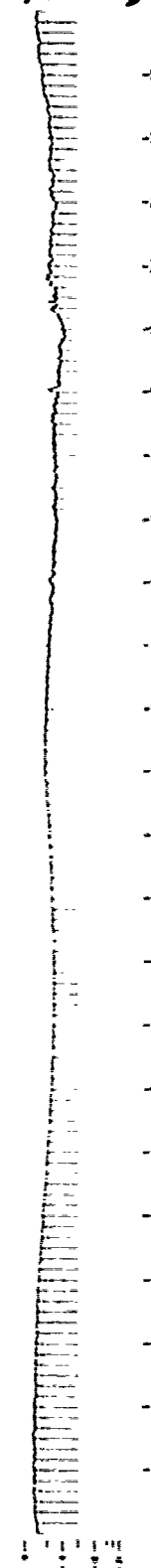
JANUARY
11
MONDAY



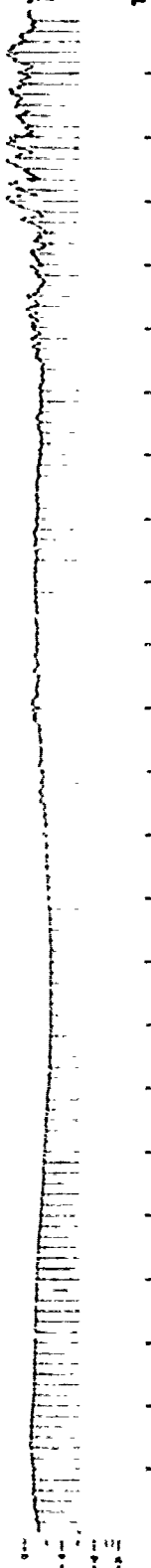
JANUARY
12
TUESDAY



JANUARY
13
WEDNESDAY



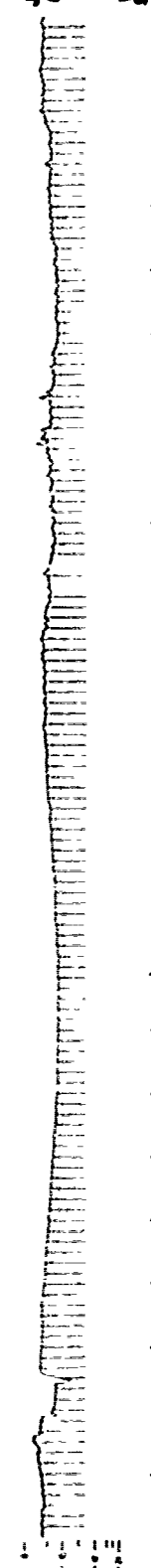
JANUARY
14
THURSDAY



JANUARY
15
FRIDAY



JANUARY
16
SATURDAY



18 MAC COSMIC NOISE INTENSITY VS UNIVERSAL TIME IN HOURS

1960

January 17 0210 - 0225 Peak, cause unknown.

January 18 Geophysical Alert Number 46: Magnetic storm started January 17 at 1200.
1805 - 1835 Peak, caused by interference.

January 19 0223 - 0242 Peak.
0320 - 0329 Peak.
0332 - 0340 Peak.
0406 - 0416 Peak.
0630 - 0640 Peak.
The peaks from 0223 to 0640 occur during an interval of strong variations of atmospheric electric potential gradient and atmospheric pressure, accompanied by precipitation.

1203 - 1211 Peak, cause unknown.

January 20 0710 - 2120 Variable.

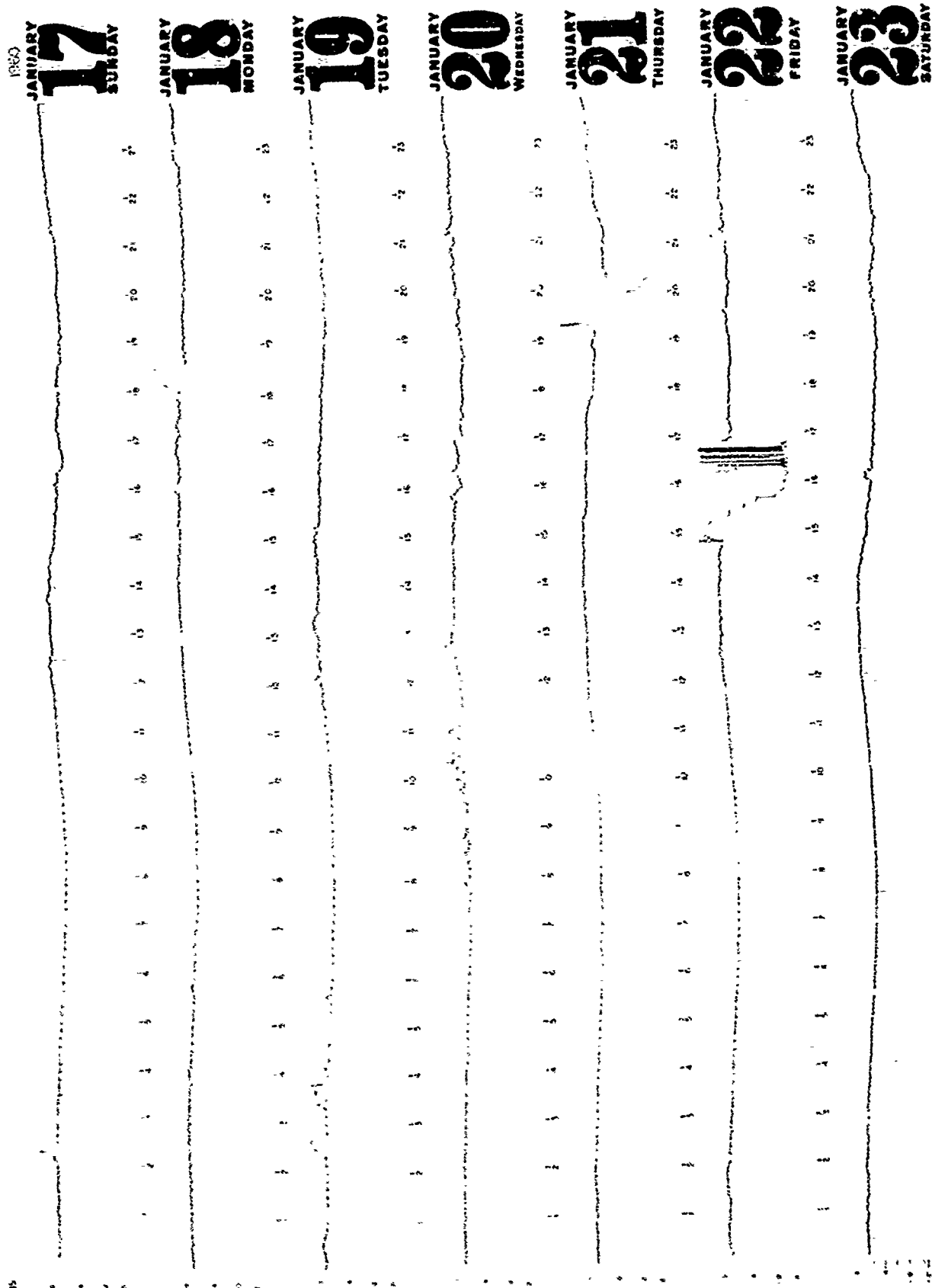
January 21 Geophysical Alert Number 47: Magnetic storm started January 21 at 00xx.
2130 - 2140 Burst.

January 22 1510 - 1620 Equipment adjustments.
2049 - 2115 Burst, correlates with dip in intensity of 27-kilocycle atmospheric.

January 23 1557 - 1612 Peak, cause unknown.

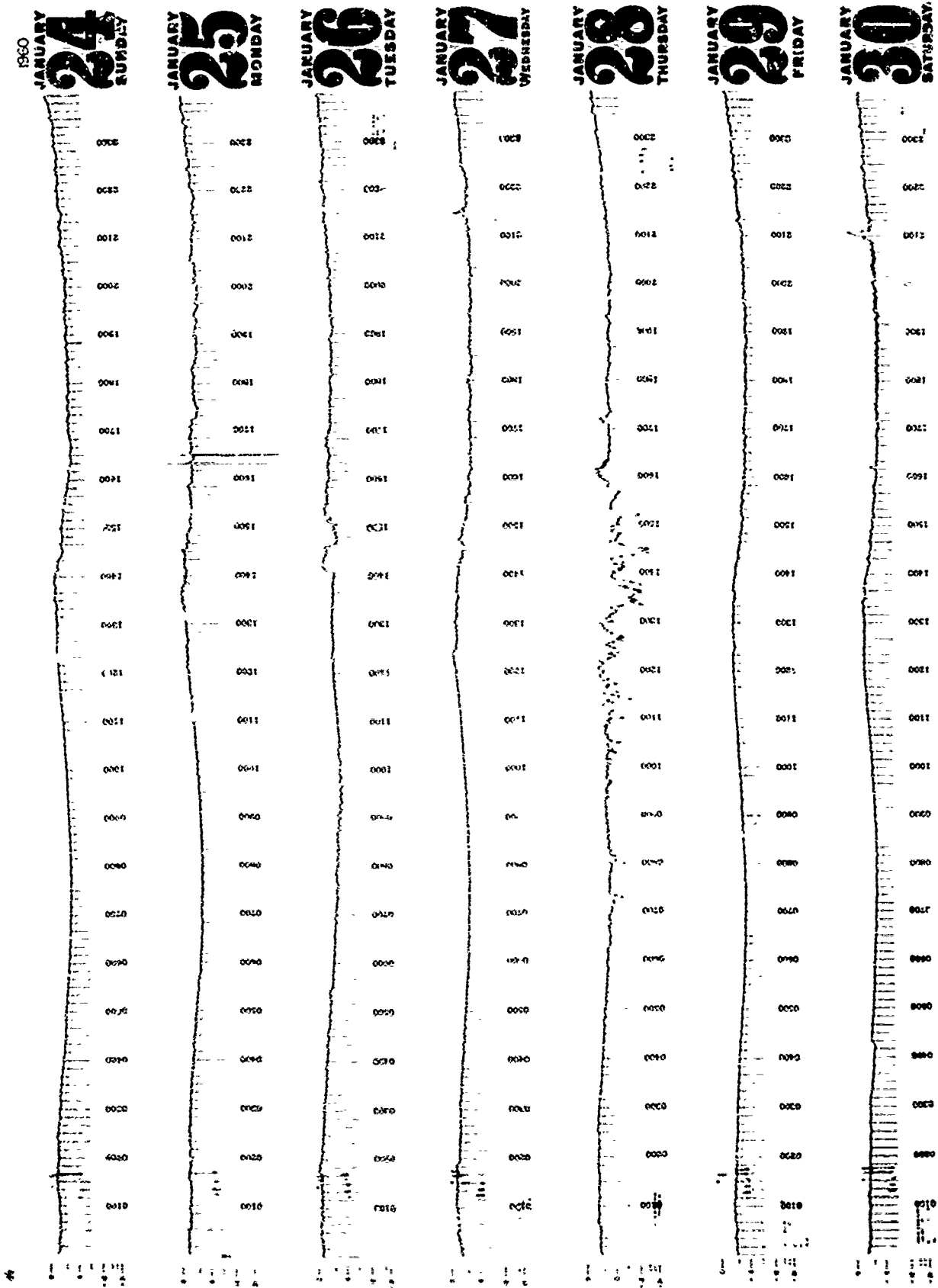
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IB MC COSMIC NOISE INTERGRY VS UNIVERSAL TIME IN HOURS

1960		
January 24	No comment.	
January 25	1620 - 1631	Equipment adjustment.
January 26	1405 - 1532	Variable.
January 27	2110 - 2125	Peak, cause unknown.
January 28	0710 - 1720	Variable, accompanied by precipitation.
January 29	No comment.	
January 30	0417 - 0420	Rise in level, cause unknown.
	1608 - 1614	Peak, cause unknown.
	2051 - 2108	Burst.



18 MC COSMIC NOISE INTENSITY VS UNIVERSAL TIME IN HOURS

1960

January 31

0810 - 1604

Chart drive not functioning properly.

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1960
JANUARY
31
SUNDAY